A test of the partitioning between ClO and ClONO₂ using simultaneous UARS measurements of ClO, NO₂, and ClONO₂

A. E. Dessler¹, S. R. Kawa¹, A. R. Douglass¹, D. B. Considine^{1,2}, J. B. Kumer³, A. E. Roche³, J. L. Mergenthaler³, J. W. Waters⁴, J. M. Russell, III⁵, J. C. Gille⁶

Abstract. The photochemical theory of the partitioning between the chlorine species ClONO₂ and ClO in the lower and mid stratosphere is tested for the first time using simultaneous measurements of ClO, NO₂, and ClONO₂. The theory suggests that over most of the day the production of ClONO₂ through the three-body reaction of ClO and NO₂ is balanced by the loss of ClONO₂ through photolysis. Our analysis tests this theory by utilizing simultaneous measurements of these species from instruments aboard the Upper Atmosphere Research Satellite (UARS), as well as calculations of the photolysis rate of ClONO₂ and laboratory measurements of the rate constant for the reaction between ClO and NO₂. Between 32 and 10 hPa (~24 and 32 km), averaged instantaneous production and loss rates of ClONO₂ agree within ~10%. At 46 hPa (~21 km), the agreement between averaged instantaneous production and loss of ClONO₂ is poorer than at higher altitudes, but still within uncertainties. This analysis provides no evidence for any problems in our understanding of the partitioning between the chlorine species ClONO₂ and ClO over the range of pressures (46 to 10 hPa) and latitudes (60°S to 60°N) considered here.

Introduction

An understanding of the effect of chlorine on the stratosphere is among the most important questions facing atmospheric scientists today. Except in the perturbed polar vortices, inorganic chlorine (Cl_v) in the lower and mid stratosphere is principally bound in the molecules HCl and ClONO₂, which together make up more than 90% of Cl_v . The ClO radical makes up a significant fraction of the remainder. How much of Cl_v is bound in each member of Cl_v is important because the ClO radical is involved in catalytic loss cycles that regulate the ozone loss over much of the stratosphere [World Meteorological Organization, 1995] while HCl and ClONO₂ are not involved in cycles that destroy ozone. An assessment of how Cl_v partitions between its member species is therefore necessary to determine the role that Cl_v plays in regulating stratospheric O_2 .

The focus in this paper is on the interaction between the species CIO and CIONO₂ between ~21 and 32 km and 60°N and 60°S. Our current understanding of the photochemistry of CIO and CIONO₂ implies that because of the

slow formation and destruction rate of HCl, the sum of mixing ratios of the species ClO and ClONO₂ in an air parcel is constant on a time scale of days to weeks. While their sum is constant, these species interconvert with a time scale of hours. During daylight, ClONO₂ is destroyed through photolysis, which most often severs the Cl-O bond [DeMore et al., 1994]. The Cl atom formed reacts rapidly (e-folding time << 1 s) with O₃ to form ClO, while the NO₃ molecule is rapidly photolyzed (e-folding time ~ 10 s) to form NO₂. We will therefore assume in this paper that ClONO₂ photolysis yields ClO and NO₂,

$$CIONO_2 + hv \longrightarrow CIO + NO_2$$
 (1)

In addition to photolysis, at high altitudes the oxidation of $CIONO_2$ by O and Cl atoms and the OH radical becomes a significant loss pathway for $CIONO_2$. At 10, 15, and 22 hPa, model estimates indicate that these radicals are responsible for ~10%, 7%, and 3% of $CIONO_2$ loss, respectively.

ClONO₂ is reformed via the three-body reaction,

$$CIO + NO_2 \xrightarrow{M} CIONO_2$$
 (2)

During most of the day, the production and loss of ClONO₂ balance and the system is in steady state [Kawa et al., 1992]. As the sun sets, photolysis of ClONO₂ slows and eventually ceases, while reaction (2) continues, causing ClO to be depleted with an e-folding time of an hour or less. At sunrise, photolysis of ClONO₂ resumes. The ClO abundance quickly rises and steady state is achieved in an hour or less.

Previous analyses have tested this theory in the lower stratosphere near 20 km altitude. Brune et al. [1990] demonstrated that the sunrise variation of ClO abundance is consistent with the photolysis of ClONO₂ in a zero-dimensional photochemical model. Extending on this,

Copyright 1996 by the American Geophysical Union.

Paper number 96JD00695. 0148-0227/96/96JD-00695\$09.00

NASA Goddard Space Flight Center, Greenbelt, Maryland.

²Applied Research Corporation, Landover, Maryland.

³Lockheed-Martin Palo Alto Research Laboratory, Palo Alto, California.

⁴Jet Propulsion Laboratory, Pasadena, California.

⁵NASA Langley Reseach Center, Hampton, Virginia.

⁶National Center for Atmospheric Research, Boulder, Colorado.

Stimpfle et al. [1994] used lower-stratosphere ClO measurements and NO2 derived from measurements of NO to demonstrate that ClO is inversely correlated with NO₂, consistent with the formation and photolysis of ClONO₂. Allen and Delitsky [1991] used measurements of CIONO₂ and NO₂ by ATMOS to infer a ClO profile. Comparisons with measurements show that the calculated ClO profile is too high. A modeling paper by Salawitch et al. [1994] concluded that lower-stratosphere measurements of ClO and NO₂ are adequately reproduced by models incorporating heterogeneous chemistry on sulfate aerosols, but that modeled HCl was a factor of two larger than measurements. To explain these and other measurements of HCl that were significantly lower than model predictions, Webster et al. [1994] proposed that overestimates by a factor of three to ten in calculation of the photolysis rate of ClONO₂ (J_{CIONO}) cause CIO in models to be overestimated relative to ClONO₂. This overprediction of ClO, coupled with the subsequent reactions,

$$ClO + NO \longrightarrow Cl + NO_{2}$$

$$Cl + CH_{4} \longrightarrow HCl + CH_{3}$$
(3)

produces a higher formation rate of HCl in models than in the atmosphere, resulting in an overprediction of HCl by models.

None of the previous analyses utilized simultaneous measurements of ClO, NO₂, and ClONO₂. In this paper we will use newly available simultaneous measurements of these species to test the theory that we have discussed above and thereby attempt to confirm the processes that control the relative abundances of ClO and ClONO₂ over the altitude range of 21 to 32 km.

Data

CIO and CIONO₂

Level 3AT constituent data from the Upper Atmosphere Research Satellite (UARS) are used in this analysis. The data are from the 46-, 32-, 22-, 15-, and 10-hPa pressure surfaces, corresponding approximately to the altitude range of 21 to 32 km. The data in our analysis extend from 60°N to 60°S and were obtained between January 1 and April 15, 1993.

Measurements of ClO were made by the Microwave Limb Sounder (MLS) at 46, 22, and 10 hPa; 32- and 15hPa abundances are obtained by interpolation. Minor adjustments have been applied to the version 3 data according to Waters et al. [1996]. The systematic uncertainty in the measurement can be a scaling error, where the error is a constant fraction of the observation, or a bias, in which case the true and measured value differ by a constant amount. Outside of the winter polar regions the bias can be as much as 0.2 parts per billion by volume (ppbv) at pressures of 46 hPa. This bias affects both day and night measurements and its magnitude can therefore be determined by examining MLS ClO measurements at night when ClO is zero. To account for the bias, we first calculate average nighttime ClO (solar zenith angles $> 100^{\circ}$) as a function of latitude for each pressure level, using all of the data obtained between January 1 and April 15, 1993. The average nighttime CIO measurement is approximately -100 parts per trillion by volume (pptv) over the range of altitudes and latitudes used in this analysis. Then, for each daytime CIO measurement used in this analysis, we subtract the average nighttime MLS CIO value at the latitude and pressure of the daytime measurement. All CIO data used in this analysis have been adjusted following this procedure.

On the basis of our comparison between MLS ClO and the difference between night and day ClONO₂ (Figure 1, discussed below) and the analysis presented by Waters et al. [1996, Figure 35], we estimate an upper limit of the bias uncertainty in the day minus night ClO data to be 50 pptv. We estimate the total accuracy of the ClO data to be the root-sum-of-squares of this 50 pptv bias uncertainty and an 8% scaling uncertainty estimated by Waters et al. [1996] (all uncertainties discussed in this paper are ~68% confidence limits, i.e., "10"). This corresponds to a total accuracy estimate ranging from $\pm 70\%$ at 46 hPa to $\pm 21\%$ at 10 hPa. Note that the precision uncertainty of any one MLS measurement is large compared to this (±300-400 pptv). Our analysis, however, averages a sufficient number of measurements to reduce the precision uncertainty to levels far below the accuracy estimate.

CIONO₂ data measured by the Cryogenic Limb Array Etalon Spectrometer (CLAES) (version 7) have a stated accuracy ranging from ±23% at 46 hPa to ±28% at 10 hPa [Mergenthaler et al., 1996]. Simultaneous measurements of both CIO and CIONO₂ allow us to perform internal consistency checks on these data to supplement the accuracy estimates provided in the literature. Because CIO is primarily produced during the day from the photolysis of CIONO₂, we have,

$$\langle \text{CIO} \rangle^{day} \cong \langle \text{CIONO}_2 \rangle^{night} - \langle \text{CIONO}_2 \rangle^{day}$$
 (4)

where the angled brackets (<>) indicate an average over the day or night, as indicated. Model simulations indicate that conversion of CIO into HOCl during the day introduces an error into equation (4) of less than 7%.

Figure 1 shows daytime average MLS ClO (left-hand side of equation (4)) versus the difference between average nighttime and average daytime CLAES ClONO2 (righthand side of equation (4)). The error bars represent the accuracy estimates; a sufficient number of measurements have been averaged in this plot so that precision uncertainty is negligible. The agreement between MLS ClO and the diurnal difference of CLAES ClONO2 is better than 50 pptv at all pressure levels. A linear fit to the data produces a line with a slope very close to one (0.95) and a small y intercept (28 pptv). At 22, 15, and 10 hPa, the difference between daytime average ClO and the diurnal difference of ClONO₂ is less than 10%, much smaller than the uncertainty. At 46 and 32 hPa, the differences are 54 and 38%, respectively, but are small in absolute terms (less than 50 pptv) and within the uncertainty estimates.

A second test of the chlorine data is obtained from an analysis of the Halogen Occultation Experiment (HALOE) HCl and nighttime CLAES ClONO₂, whose sum closely approximates Cl_y at the altitudes being considered. Figure 1 of *Dessler et al.* [1995] (not shown) compares the sum of HCl and nighttime ClONO₂ from UARS with Cl_y derived from a relation based on aircraft data [Woodbridge et al.,

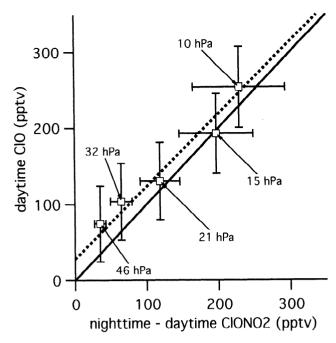


Figure 1. Daytime average Microwave Limb Sounder (MLS) CIO (pptv) versus nighttime average Cryogenic Limb Array Etalon Spectrometer (CLAES) ClONO₂ minus daytime average CLAES ClONO₂ (pptv). Daytime is defined as solar zenith angles less than 80°, nighttime as solar zenith angles greater than 100°. The solid line is the one-to-one line. The dotted line is a least squares fit to the UARS comparison (MLS=0.95*CLAES+28 pptv). For each pressure level, daytime average ClO and ClONO2 are computed for 10°-wide latitude bins using simultaneously measured ClO and ClONO₂. Nighttime ClONO₂ is also averaged in 10°-wide latitude bins. Then, the daytime average ClONO2 concentration is subtracted from the nighttime ClONO₂ average at the same latitude. The latitude bins are then averaged to produce a single average ClO and average nighttime minus daytime ClONO₂ for the entire pressure surface. The error bars represent the accuracy estimates and are the root-sum-of-squares of an 8% scaling uncertainty and a 50 pptv bias uncertainty for ClO and a 23-28% scaling uncertainty for nighttime minus daytime ClONO₂. Data were obtained between January 1 and April 15, 1993, and cover the latitude range between 60°N and 60°S.

1995]. The agreement between Cl_y and the sum of HCl and nighttime ClONO₂ between potential temperatures of 550 and 800 K (~46 and 10 hPa) provides supportive evidence that the ClONO₂ and HCl measurements are accurate. Note that these internal consistency checks use measurements from different instruments utilizing different measurement techniques and different retrieval algorithms to perform simple conservation tests on the data set. The ability to perform such tests as well as the good agreement found underscores the usefulness of the UARS data set.

NO_2

Measurements of NO₂ by the CLAES are made simultaneously with the ClO and ClONO₂ data. However, these version 7 CLAES NO₂ data have not been validated at the pressures used in this analysis. To address this issue, we

use NO and NO₂ measurements (both version 17) from the HALOE to adjust the CLAES NO₂ data. HALOE is a solar occultation instrument and it therefore makes all of its measurements at solar zenith angles of 90°. At this solar zenith angle, NO and NO₂ are rapidly interconverting and it is most convenient to deal with the sum of NO and NO₂, which is traditionally denoted as NO_x.

We first compare the CLAES NO_2 data with the HALOE NO_x data. Because NO is converted to NO_2 at night with a time scale of minutes, NO_x at night is constituted almost entirely of NO_2 . Therefore HALOE sunset NO_x is comparable to CLAES NO_2 measured just after sunset; HALOE sunrise NO_x is comparable to CLAES NO_2 measured just before sunrise. Figure 2 compares the ratio of near-sunset to near-sunrise NO_2 from CLAES with ratio of sunset to sunrise NO_x in HALOE. One complication of this comparison is that NO_x decreases through the night because of the reaction

$$NO_2 + O_3 \longrightarrow NO_3 + O_2 \tag{5}$$

Thus, NO_2 measured just after sunset will be slightly smaller than sunset NO_x , and NO_2 measured just before

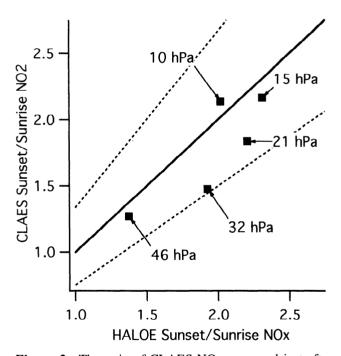


Figure 2. The ratio of CLAES NO₂ measured just after sunset (averaged over solar zenith angles between 95° and 110°) to CLAES NO₂ averaged over those same solar zenith angles just before sunrise versus the ratio of HALOE sunset NO_x to sunrise NO_x. The CLAES data have been adjusted for the small change in NO₂ between the measurement time and sunset/sunrise due to the reaction with ozone (see text). For each pressure level, zonally averaged CLAES and HALOE ratios are computed for 10°wide latitude bins. These are then averaged to produce one value for each pressure surface, and these average values are plotted as squares. The solid line is the one-to-one line. The dashed lines are $\pm 25\%$. Data in these plots were obtained between January 1 and April 15, 1993, and between 60°N and 60°S. Because a small offset in the NO₂ measurements at low concentrations could result in a large percent error, we have not included data between 20°N and 20°S on the 46, 32, and 22-hPa surfaces.

sunrise will be slightly larger than sunrise NO_x . We account for this effect by adjusting the near-sunset/sunrise CLAES NO_2 data by $2k_{NO2+O3}[O_3]\Delta t$, the fraction of NO_x depleted through reaction (5) in a length of time Δt . The factor of two arises because each NO_3 formed via equation (5) quickly reacts with another NO_2 to form N_2O_5 . We calculate this correction factor using zonally-averaged monthly-mean O_3 and temperature, both measured by MLS (version 3). This adjustment increases the diurnal difference in the CLAES data by ~10%. Figure 2 shows general agreement between the magnitudes of the diurnal cycles in the HALOE NO_x and CLAES NO_2 data. The agreement is within 10% at 10, 15, and 46 hPa and within 17% and 24% at 22 and 32 hPa, respectively.

While the relative magnitude of the diurnal cycle is similar in the two data sets, comparisons of the absolute magnitude of the CLAES NO₂ and the HALOE NO_x show differences of as much as a factor of two. This suggests that these CLAES data are internally consistent but have an absolute calibration error. To correct the CLAES NO2 data, we first determine the ratio of zonally averaged HALOE NO_x to zonally averaged CLAES NO₂ at sunrise and sunset as a function of latitude and pressure. Note that the CLAES NO2 measurements are taken from just before sunrise and just after sunset and adjusted to sunrise and sunset as discussed previously. Then, each daytime CLAES NO₂ measurement is multiplied by the average of the sunrise and sunset ratios at the measurement's latitude and pressure. In other words, we are creating an NO2 data set that uses the information from the CLAES about the relative changes in NO₂ throughout the day and information about the absolute calibration information from the HALOE. We will hereinafter refer to the CLAES NO₂ data that has been adjusted by the HALOE data as the "adjusted NO2 data".

The accuracy of both the HALOE NO and NO₂ measurements is estimated to be $\pm 15\%$ [Gordley et al., 1996], leading to a root-sum-of-squares of the uncertainty in NO_x of $\pm 21\%$. From Figure 2 we estimate the uncertainty in the diurnal cycle of the CLAES NO₂ data to be $\pm 25\%$. Combining these uncertainties, we calculate that the overall uncertainty in the adjusted NO₂ data is $\pm 33\%$. Figure 3 is a plot of diurnal cycle of the adjusted NO₂ data. The shape of the profile is in good agreement with that seen by Webster et al. [1990]. Also shown is an average of the unadjusted version 7 CLAES NO₂ data and the HALOE NO_x data used to adjust the CLAES NO₂ data.

Ancillary Data

Temperature measurements by the CLAES are used in this analysis [Gille et al., 1996]. The accuracy of these measurements is estimated to be ±2 K. Comparisons between the CLAES data with temperatures from the National Meteorological Center (NMC) Climate Analysis Center [Nagatani et al., 1990 and references therein] interpolated to the location of the CLAES measurements reveal average disagreements smaller than 1% at both 10 and 46 hPa. The rms disagreement is 3% and 1% at 10 and 46 hPa, respectively. Temperatures are used to calculate the ambient number density, the rate constant for the three-body reaction between ClO and NO₂, and the cross section of ClONO₂. The uncertainty in the temperature is negli-

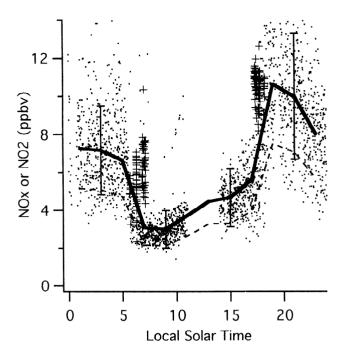


Figure 3. NO₂ and NO_x versus local solar time at 10 hPa. Individual adjusted NO₂ data (dots), the average of the adjusted data (solid line), and average of unadjusted CLAES NO₂ data (dashed line) are shown. The error bars on the adjusted NO₂ average line are $\pm 33\%$ and represent the accuracy of the data. Also shown are HALOE NO_x measurements (+) at sunrise and sunset. CLAES data were obtained between February 12 and March 16, 1993, and between 25°N and 35°N. HALOE data were obtained between January 20 and April 10, 1993, and between 25°N and 35°N.

gible compared to the other sources of uncertainty in this analysis and will not be considered further.

The rate constant for the three-body reaction between ClO and NO₂ (reaction (2)) is taken from the JPL 94 assessment of *DeMore et al.* [1994]. Using information in this source we estimate the 1σ error is ~ $\pm 30\%$. The photolysis rate of ClONO₂, J_{ClONO_2} , is calculated as a function of pressure, temperature, overhead O₃, and solar zenith angle. The photolysis rate routine is taken from the Goddard three-dimensional chemical transport model [Kawa et al., 1995] and is based on radiative transfer calculations from the model of Anderson and Lloyd [1990]. The ClONO2 cross section is from the JPL 94 assessment of DeMore et al. [1994] and is temperature dependent. Overhead ozone is determined from a zonally averaged, monthly climatology of MLS O₃ measurements (205 GHz, version 3) obtained between January 1 and April 30, 1993. The radiation calculations are for clear-sky conditions and a surface albedo of 0.3. We estimate the uncertainty in the cross section of ClONO₂ to be ±15% [Burkholder et al., 1994]. We estimate uncertainties in the radiative calculation to also be ±15%, leading to a combined 1 σ uncertainty for J_{ClONO_2} of $\pm 21\%$.

Balance Between Production and Loss of ClONO2

To investigate our understanding of the ClO - ClONO₂ system, we compare the instantaneous production and loss

rate of ClONO₂. Loss of ClONO₂ is primarily through equation (1), photolysis of ClONO₂, and the rate of this process is $J_{\text{ClONO}_2}|\text{ClONO}_2|$. Production of ClONO₂ is through equation (2), the three-body reaction of ClO and NO₂, and the rate of this process is $k[\text{ClO}][\text{NO}_2][\text{M}]$. Figure 4a shows this comparison on the 22-hPa surface. There are approximately ten thousand points in this plot; each one represents a simultaneous measurement of ClO, NO₂, and ClONO₂ and a calculation of the J_{ClONO_2} and k. The data show considerable scatter, forming a cloud around the one-to-one line. The origin of the scatter in Figure 4a is not variability in the atmosphere but is predominantly due to the large precision uncertainty of the ClO data, the statistics of which have been discussed by *Schoeherl et al.* [1993] and *Waters et al.* [1996].

To eliminate the precision uncertainty, we average the instantaneous production and loss rates. The average production and loss rate for this pressure surface appears in Figure 4a as the square. Note that the average falls very close to the one-to-one line, indicating that production and loss of ClONO₂ balance. Figure 4b shows a close-up of the averaged value, graphically displaying the agreement between averaged instantaneous production and loss of ClONO₂. Also shown are the accuracy error bars, obtained from propagation of the uncertainties listed in Table 1.

Despite the large spread in the raw data, the precision uncertainty of the average is only a few percent, insignificant compared to the accuracy uncertainty.

Figure 5 compares the averaged instantaneous production and loss rates of ClONO2 for the five pressure levels considered. For each pressure level, the individual measurements have been averaged, as discussed in the last paragraph, and only the averaged value for each pressure level is plotted. The error bars again represent the accuracy estimate of the average. In general the agreement is quite good. At 32 and 22 hPa, production and loss agree to within 1%. At 15 and 10 hPa, production exceeds loss by 13 and 23%, respectively. Part of this discrepancy results from our neglect of the loss of ClONO₂ at high altitudes through reactions with O, Cl, and OH. We estimate that these oxidation pathways contribute 7 and 10% to the loss of ClONO₂ at 15 and 10 hPa, respectively. Considering these neglected oxidation processes, the agreement between the averaged production and loss rates at 15 and 10 hPa is ~6 and 13%, respectively. At 46 hPa, production exceeds loss by about 50%. While within uncertainty limits, this agreement is much poorer than at higher altitudes.

We conclude from our analysis that there is no evidence to support missing chemistry, a significant error in the rate constant for the reaction of CIO and NO₂, or errors in our

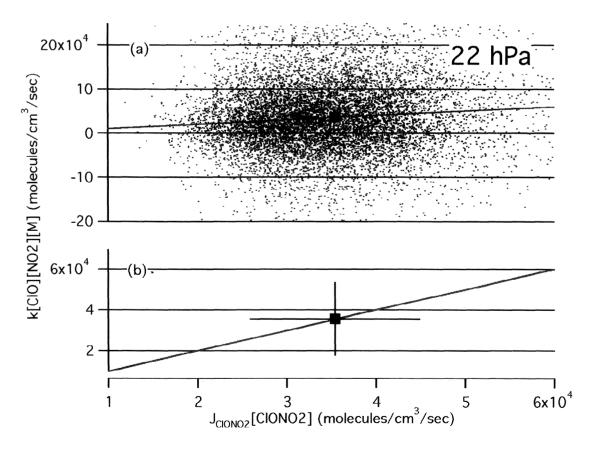


Figure 4. (a) Instantaneous production of CIONO₂ (k[CIO][NO₂][M]) versus loss of CIONO₂ (J_{CIONO_2} [CIONO₂]) on the 22-hPa pressure surface. Each dot represents a simultaneous measurement of CIO, NO₂, and CIONO₂ and a calculation of J_{CIONO_2} and k. The square is the average production and loss rate. The line is the one-to-one line. (b) Close-up of average production and loss rate. Error bars represent accuracy estimates; precision uncertainty is negligible and not considered. Note the agreement between average production and loss. The data in both Figures 4a and 4b were obtained between January 1 and April 15, 1993, and between 20° and 60° latitude in both hemispheres.

Table 1. Errors

Table 1. Liivis		
Measurement	Accuracy, %	_
CIONO ₂	±23-28	
ClO	±21-70	
Adjusted NO ₂	±33	
J	±21	
k	±30	
J[CIONO ₂]	±31-35	
$k[ClO][NO_2][M]$	±49-83	

These uncertainties refer to ~68% confidence limits; that is, they are " 1σ ".

calculation of the photolysis rate of ClONO₂. In particular, we see no evidence to support previous suggestions of large pressure-dependent errors in calculations of the photolysis rate of ClONO₂ [Webster et al., 1994]. Note that the 46-hPa data show production exceeding loss. This is opposite to the effect suggested by Webster et al. [1994], who would predict loss of ClONO₂ exceeding production when using standard J_{ClONO_2} calculations. Our analysis of the raw 46-hPa data suggests that the degraded agreement between production and loss at 46 hPa is a result of uncertainty in the UARS data at this altitude and is not likely related to problems with our theoretical understanding of the ClO - ClONO₂ system.

We have shown in this figure that the production and loss rates of ClONO₂ balance on average. Because of the precision uncertainty in the data we cannot demonstrate balance on a point-by-point basis. However, while we cannot rule out the possibility that production exceeds loss in some region of the atmosphere and loss exceeds production in another region of the atmosphere such that the imbalances cancel each other, there is no a priori reason to believe that an imbalance in one region of the atmosphere would be canceled to within a few percent by an opposite imbalance in some other region of the atmosphere. We conclude therefore that it is unlikely for ClONO₂ production and loss rates to deviate significantly from balance on a point-by-point basis, although this point deserves further research.

A natural question arising from this analysis is how big a missing source or sink of $ClONO_2$ would be consistent with our analysis. Because our estimate of the production and loss rates of $ClONO_2$ balance to within ~10% between 10 and 32 hPa, any significant missing source or sink of $ClONO_2$ must fortuitously cancel (to within ~10%) an opposite systematic error in the data over the entire altitude extent of the missing chemistry. If we assume that the missing chemistry affects only one of the pressure levels in our data set, then we cannot rule missing chemistry that

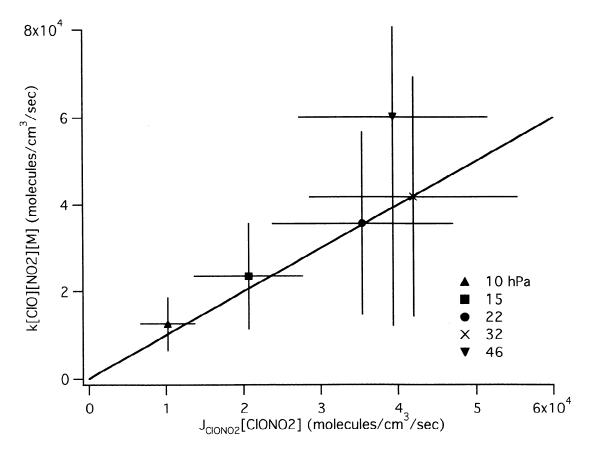


Figure 5. Production of ClONO₂ (k[ClO][NO₂][M]) versus loss of ClONO₂ (J_{ClONO₂}[ClONO₂]) for the five pressure levels analyzed in this paper. The data have been binned and averaged as discussed in text. Error bars represent accuracy estimates; precision uncertainty is negligible and not considered. The solid line is the one-to-one line. Note the agreement between production and loss. The data were obtained between January 1 and April 15, 1993. The 15 and 10-hPa data cover the latitude range between 60°N and 60°S. The 46-, 32-, and 22-hPa data cover the latitude range between 20° and 60° in each hemisphere.

is smaller than the error bars shown in Figure 5 for that level. However, missing chemistry that affects only one of our pressure levels is unlikely and puts sharp constraints on the potential mechanisms. Missing chemistry with a significant vertical extent (e.g., affecting several pressure levels) would require canceling systematic errors in the data on several pressure levels. Based on analyses of the retrieval algorithms, we estimate that the systematic errors for each given pressure level are largely uncorrelated with the systematic errors on other pressure levels. We believe therefore that any missing chemistry affecting several pressure levels must be significantly smaller than the error bars on any one level indicate. Without information about the altitude dependence of the missing chemistry, however, it is impossible to quantitatively evaluate the probability of the existence of the missing chemistry.

Due to the constraint of steady state, our analysis has been limited to investigating solar zenith angles of less than 80°. We have seen no evidence to suggest problems with our understanding of the production and loss of ClONO2 at solar zenith angles less than 80°. However, the partitioning of Cl_y between HCl and ClONO₂ has a time constant that is weeks to months in the lower and mid stratosphere and is therefore a function of the production and loss rates of ClONO2 averaged over all daytime solar zenith angles, including those greater than 80°. Calculations using the radiative model, however, indicate that solar zenith angles greater than 80° contribute only ~10 to 20% to the daytime-averaged production and loss of ClONO₂ over the latitude range of 60°N to 60°S. In conjunction with our analysis, we therefore conclude that the daytime-averaged production and loss of CIONO2 are similarly well understood (within our uncertainties) and are unlikely to be significantly in error.

Conclusions

We have tested the partitioning between the chlorine species $ClONO_2$ and ClO using simultaneous measurements of ClO, NO_2 , and $ClONO_2$ for the first time. We conclude the following:

- 1. Between 32 and 10 hPa (~24 and 32 km), average production and loss rates of ClONO₂ agree within a few percent, as expected by photochemical theory. At 46 hPa (~21 km), the agreement between averaged production and loss rates of ClONO₂ is poorer than at higher altitudes but still within uncertainties.
- 2. We see no evidence for any problems in our understanding of the partitioning between the chlorine species ClONO₂ and ClO over the range of pressures (46 to 10 hPa) and latitudes (60°S to 60°N) considered here. We conclude that our understanding of the chemical mechanisms, photolysis rate of ClONO₂, and the rate constant for the reaction between ClO and NO₂ is good within the uncertainties of this analysis.

Acknowledgments. This work was supported by grants from the NASA Atmospheric Chemistry Modeling and Analysis Project and the UARS Guest Investigator Program. A.E.D. was supported by a National Research Council postdoctoral fellowship. We acknowledge discussions and suggestions from R. C. Cohen, L. Froidevaux, E. J. Hintsa, C. H. Jackman, L. J. Mickley, M. R. Schoeberl, R. S. Stolarski, and P. O. Wennberg. NMC data were provided by L. R. Lait, P. A. Newman, and M. R. Schoeberl.

References

- Allen, M., and M. L. Delitsky, Inferring the Abundances of CIO and HO₂ From Spacelab 3 Atmospheric Trace Molecules Spectroscopy Observations, J. Geophys. Res., 96, 2913-2919, 1991.
- Anderson, D. E., Jr., and S. A. Lloyd, Polar twilight UV-visible radiation field: Perturbations due to multiple scattering, ozone depletion, stratospheric clouds, and surface albedo, J. Geophys. Res., 95, 7429-7434, 1990.
- Brune, W. H., D. W. Toohey, S. A. Lloyd, and J. G. Anderson, The sunrise and sunset variation of ClO in the lower stratosphere, *Geophys. Res. Lett.*, 17, 509-512, 1990.
- Burkholder, J. B., R. K. Talukdar, and A. R. Ravishankara, Temperature dependence of the ClONO₂ UV absorption spectrum, *Geophys. Res.* Lett., 21, 585-588, 1994.
- DeMore, W. B., et al., Chemical kinetics and photochemical data for use in stratospheric modeling, JPL Pub. 94-26, Jet Propulsion Laboratory, Pasadena, Calif., 1994.
- Dessler, A. E., et al., Correlated observations of HCl and ClONO₂ from UARS and implications for stratospheric chlorine partitioning, *Geophys. Res. Lett.*, 22, 1721-1724, 1995.
- Gille, J. C., et al., Accuracy and precision of Cryogenic Limb Array Etalon Spectrometer (CLAES) temperature retrievals, J. Geophys. Res., in press, 1996.
- Gordley, L. L., et al., Validation of nitric oxide and nitrogen dioxide measurements made by HALOE for the UARS platform, J. Geophys. Res., in press, 1996.
- Kawa, S. R., et al., Photochemical partitioning of the reactive nitrogen and chlorine reservoirs in the high-latitude stratosphere, J. Geophys. Res., 97, 7905-7923, 1992.
- Kawa, S. R., et al., Missing chemistry of reactive nitrogen in the upper stratospheric polar winter, Geophys. Res. Lett., 22, 2629-2632, 1995.
- Mergenthaler, J. L., et al., Validation of CLAES CIONO₂ measurements, J. Geophys. Res., in press, 1996.
- Nagatani, R. M., A. J. Miller, M. E. Gelman, and P. A. Newman, A comparison of Arctic lower stratospheric winter temperatures for 1988-89 with temperatures since 1964, *Geophys. Res. Lett.*, 17, 333-336, 1990.
- Salawitch, R. J., et al., The distribution of hydrogen, nitrogen, and chlorine radicals in the lower stratosphere: Implications for changes in O₃ due to emission of NO_y from supersonic aircraft, *Geophys. Res. Lett.*, 21, 2547-2550, 1994.
- Schoeberl, M. R., et al., MLS CIO observations and Arctic polar vortex temperatures, Geophys. Res. Lett., 20, 2861-2864, 1993.
- Stimpfle, R. M., et al., The response of CIO radical concentrations to variations in NO₂ radical concentrations in the lower stratosphere, *Geophys. Res. Lett.*, 21, 2543-2546, 1994.
- Waters, J. W., et al., Validation of UARS MLS CIO measurements, J. Geophys. Res., in press, 1996.
- Webster, C. R., et al., Hydrochloric acid and the chlorine budget of the lower stratosphere, *Geophys. Res. Lett.*, 21, 2575-2578, 1994.
- Webster, C. R., R. D. May, R. Toumi, and J. A. Pyle, Active nitrogen partitioning and the nighttime formation of N₂O₅ in the stratosphere: Simultaneous in situ measurements of NO, NO₂, HNO₃, O₃, and N₂O using the BLISS diode laser spectrometer, *J. Geophys. Res.*, 95, 13,851-13,866, 1990.
- Woodbridge, E. L., et al., Estimates of total organic and inorganic chlorine in the lower stratosphere from in situ and flask measurements during AASE II. J. Geophys. Res., 100, 3057-3064, 1995.
- World Meteorological Organization, Scientific assessment of ozone depletion: 1994, Global Ozone Res. and Monit. Proj., WMO Rep. 37, Geneva, 1995.
- D. B. Considine, A. E. Dessler, A. R. Douglass, and S. R. Kawa, NASA Goddard Space Flight Center, Code 916, Greenbelt, MD 20771 (e-mail: dessler@maia.gsfc.nasa.gov)
- J. B. Kumer, J. L. Mergenthaler, A. E. Roche, Lockheed-Martin Palo Alto Research Laboratory, Code 9120/Building 252, 3251 Hanover Street, Palo Alto, CA 94304-1187
- J. M. Russell, III, NASA Langley Research Center, MS 401B, Hampton, VA 23681-0001
- J. W. Waters, Jet Propulsion Laboratory, Mail Code 183-701, 4800 Oak Grove Drive, Pasadena, CA 91109-8099
- J. C. Gille, National Center for Atmospheric Research, Atmospheric Chemistry Division. 1850 Table Mesa Drive, P. O. Box 3000, Boulder, CO 80307-3000

(Received August 11, 1995; revised January 31, 1996; accepted January 31, 1996)